

Climate

The role of the natural atmosphere in the climate system

Human impact on the atmosphere and the climate

- Radiation balance
- The greenhouse effect
- Radiative forcing
- Climate change

Literature connected with today's lecture:

Jacob, chapter 7

Exercises:

7:1 – 7:6

Radiation

- Total radiation from a black body:

$$F = \sigma T^4; \quad \sigma = \text{Stefan-Boltzmann constant}$$

Unit of F: [W m⁻²]

- Planck's law (black body):
$$\frac{dF}{d\lambda} = \frac{2\pi hc^2}{\lambda^5 (\exp(\frac{hc}{kT\lambda}) - 1)}$$

- Kirchhoff's law:

An object (e.g. the atmosphere) absorbs fraction ε [0 - 1] at wavelength $\lambda \Rightarrow$ emission of the same fraction of the black body radiation for the temperature of the body and at that wavelength

Exercise 7:1

7:1) The sun with radius $r = 6.96 \cdot 10^8$ m emits radiation like a black body of $T = 5780$ K.
Calculate the total power [effekt] emitted by the sun!

Radiation intensity from a blackbody :

$$F = \sigma T^4 = 5.67 \cdot 10^{-8} \cdot 5780^4 \text{ W/m}^2 = 6.33 \cdot 10^7 \text{ W/m}^2$$

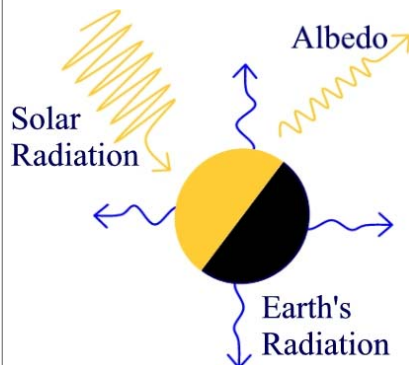
The power irradiated from the sun:

$$P = F \cdot A = F \cdot 4\pi r^2 = 4\pi \cdot (6.96 \cdot 10^8)^2 \cdot 6.33 \cdot 10^7 \text{ W} = 3.85 \cdot 10^{26} \text{ W}$$

Radiation Balance of the Earth

- Solar radiation
 - Approximately black body at 5800 K
- \Rightarrow Radiation to the Earth (solar constant):

$$F_s = \frac{P}{4\pi d^2} = \frac{4\pi R_s^2 \sigma T_s^4}{4\pi d^2} = 1370 \text{ W/m}^2$$



- Satellites show that the earth **reflects** 28% of the solar radiation (albedo; $A = 0.28$)

- Reflected: $\Phi = F_s A$
- Absorbed by the earth: $\Phi = (1 - A) F_s$

- **Absorption** by the earth (average per area)
 $\Phi = F_s(1-A)\pi R_E^2 / 4 \quad \pi R_E^2 = F_s(1-A)/4$

- Radiation balance

- The earth emits the same amount
- Consider the earth as a black body:

$$\sigma T_E^4 = F_s(1-A)/4; \quad T_E = 257 \text{ K}$$

- The earth viewed from space
 - **effective temperature** -16°C

- Climate – Temperature at the surface

Exercise 7:2

b: Planet H is at distance $d_H = 8 \cdot 10^7$ km from sun with the albedo $A_H = 0.85$. The solar constant at the earth $F_{SE} = 1370$ W/m² with $d_E = 1.5 \cdot 10^8$ km and $A_E = 0.28$. Calculate the effective temperature of the two planets!

Effective temperature:

Radiation balance $\Rightarrow (1-A)F_S/4 = \sigma T^4 \Rightarrow$

$$T = [(1-A) F_S/4\sigma]^{0.25}$$

Earth:

$$T_E = [(1-A_E)F_{SE}/4\sigma]^{0.25} = 257 \text{ K}$$

Planet H:

We need the solar constant: $F_S \sim 1/d^2 \Rightarrow$

$$F_{SH}/F_{SE} = d_E^2/d_H^2 \Rightarrow$$

$$F_{SH} = F_{SE} \cdot d_E^2/d_H^2 = 4820 \text{ W/m}^2 \Rightarrow$$

$$T_H = [(1-A_H)F_{SH}/4\sigma]^{0.25} = 238 \text{ K}$$

c: Based on the properties of the two planets: Can we say anything about the climate on the surface of the two planets?

The effective temperature tells us how much a planet radiates to space

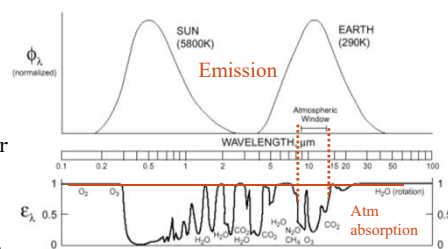
Part of that radiation from the atmosphere

The atmosphere also affects the radiation at the surface and hence the climate

ANSWER: NO

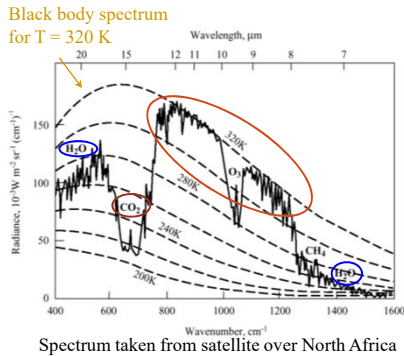
Atmospheric Absorption of Radiation

- The sun – solar radiation
 - High temperature – short wavelength
- The earth – terrestrial radiation
 - Lower temp. – longer wavelength
- UV absorbed by O₂ and O₃
- The atmosphere fairly transparent for solar radiation (H₂O)
- H₂O, CO₂, CH₄, N₂O, O₃ and CFC absorbs large fraction of the terrestrial radiation
- Atmospheric window ($\lambda = 8 - 13 \mu\text{m}$): Radiation direct from the surface to space



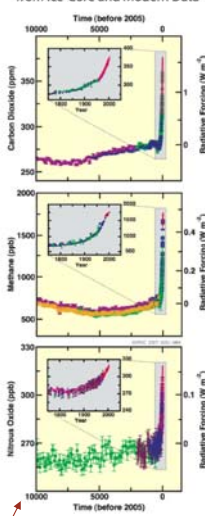
Radiation Spectrum of the Earth

- The earth radiation spectrum differs from black body radiation
 - Absorption and re-emission in the atmosphere
 - The atm temp decreases with altitude
 - Temp determines radiation (σT^4) – Less radiation from a cold body
- Combination of black bodies of different temperatures
- Atmospheric “window” (8-13 μm):
 - Weak absorption in the atm \Rightarrow
 - Radiation from **surface $T \approx 320\text{K}$** (North Africa!)
- CO_2 at 15 μm :
 - $T \approx 215\text{K}$
 - effective emission altitude ~ 10 km
- H_2O (7 & 20 μm):
 - $T \approx 260\text{K}$
 - effective emission altitude ~ 5 km (precipitation keeps H_2O at low altitudes)
- The greenhouse effect:
 - Atm absorption + atm temp decrease with altitude \Rightarrow
 - Part of the terrestrial emission from lower temp (“deep valleys” in the spectrum) \Rightarrow
 - **Radiation balance**
 - Total emission corresponding to 257 K (T_E) needed
 - \Rightarrow increased overall temperature to compensate for the dips in the spectrum
 - \Rightarrow Increased **surface temperature**



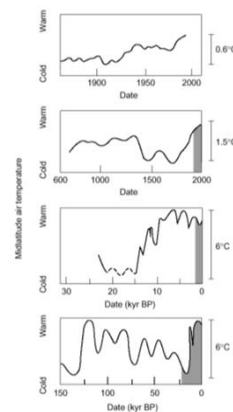
Climate Change Caused by Man?

Changes in Greenhouse Gases from Ice-Core and Modern Data



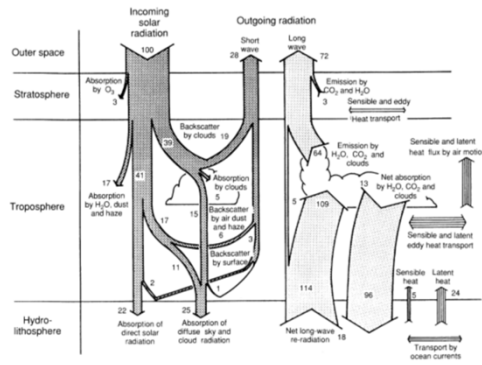
Last ice age, 11 500 y ago

- Natural GH effect (CO_2 , H_2O) increases surface temp by 33°C
- Ongoing climate change:
 - Increasing air and ocean temp
 - Reduced snow cover
 - Rising sea level
- Increased GHG concentration has increased the GH effect
- GHG emissions explain climate change?
 - Historical temp variations larger – Due to variation in solar activity (?)
- Natural variations explain the increased temperature?
 - Contradicted by climate models, which predict further warming – IPCC (UN climate panel):
 - “ It is extremely likely that human influence has been the dominant cause of observed warming since the mid 20th century.”



Radiation Balance

Global energy balance (annual mean)



Short wave radiation (in):

- 28% reflected (19% clouds)
- 72% absorbed (47% surface)

Long wave radiation:

- 96% atmosphere → surface

Non-radiative components:

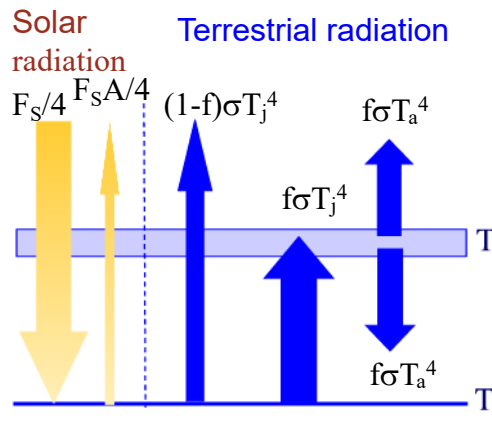
- 24 + 5% heat transport from surface to atmosphere (vapour formation, convection)

The greenhouse effect:

- The balance of the surface
 - $\Phi_{rel} = 22 + 25 + 96 - 24 - 5 = 114\%$
- The greenhouse effect increases the radiation to the surface
 - 14% stronger than the total radiation from the sun

A Simple Climate Model

- Assumptions on the atmosphere:**
 - Thin, isothermal layer
 - Absorbs fraction f of the terrestrial radiation
 - Transparent to solar radiation



Radiation balance:

Earth + atmosphere:

$$\frac{F_s(1-A)}{4} = (1-f)\sigma T_j^4 + f\sigma T_a^4$$

The Atmosphere: $f\sigma T_j^4 = 2f\sigma T_a^4$

⇒

$$T_j = \left(\frac{F_s(1-A)}{4\sigma(1-f/2)} \right)^{1/4}$$

Average surface temperature of the Earth: 288 K
- Obtained with $f = 0.77$

Exercise 7:3c

c: Planet H at $8 \cdot 10^7$ km from sun with albedo 0.85 has an atmosphere that absorbs 40% of the longwave radiation. How much does its atmospheric greenhouse effect change the temperature at the surface of planet H?

Without atmosphere:

The average surface temperature equals the effective temperature

$$T_{\text{surface1}} = T_{\text{eff}} = [(1-A_H)F_{\text{SH}}/4\sigma]^{0.25} = 237.6 \text{ K}$$

With an atmosphere (according to the model):

$$T_{\text{surface2}} = [(1-A_H)F_{\text{SH}}/(4\sigma(1-f_H/2))]^{0.25}$$

From previous exercise we know $F_{\text{SH}} = 4820 \text{ W/m}^2$

$$\text{Enter numbers } T_{\text{surface2}} = 251.2 \text{ K}$$

$$\Delta T = T_{\text{surface2}} - T_{\text{surface1}} = 13.6 \text{ K}$$

Climate Change

- Change of the radiation properties causes:
 - Initial temperature change
 - Feedbacks due to change
- Combines to a climate change

H₂O feedbacks

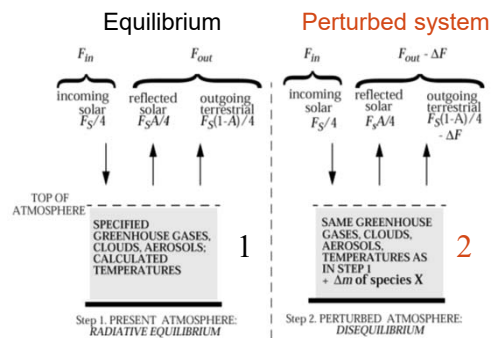
- H₂O the most important greenhouse gas (GHG)
 - Man's emissions small compared with the natural sources
- Increase of another GHG \Rightarrow increased temp \Rightarrow evaporation of H₂O \Rightarrow further increased temp \Rightarrow more evaporation...
- Counteraction
 - cloud formation and precipitation
 - Prevents H₂O from reaching high (cold) altitudes
- The role of clouds in a temperature change unclear:
 - more H₂O \Rightarrow increased cloudiness \Rightarrow increased albedo
 - more H₂O \Rightarrow larger cloud drops \Rightarrow faster formation of precipitation \Rightarrow reduced cloudiness \Rightarrow reduced albedo
- Large quantitative uncertainties concerning clouds in the climate system

Climate Change

- Feedbacks due to change
 - complicated
 - large quantitative uncertainties
- The initial phase is directly connected with the radiative properties
 - better understood quantitatively
- The potential of climate change of e.g. greenhouse gases known with high accuracy
- **Radiative Forcing** (“strålningsdrivning”)

Radiative Forcing

- Which is the initial change caused by a change in radiative properties?
- Starting conditions (Equilibrium)
 - Radiation model
 - Specify system components
 - Compute temperatures
- Perturb the system
 - Specify new components
 - Temperatures are kept unchanged
 - Causes difference between incoming and outgoing radiation
 - **Radiative Forcing** ΔF [W/m^2]
- ΔF a “theoretical” product (nature does not “freeze” the starting temperatures)
- ΔF frequently used to describe the potential of climate perturbations



$$\Delta F_1 = 0; \quad \Delta F_2 = F_{in,2} - F_{ut,2}$$

Exercise 7:3d

The earth's atmospheric concentration of greenhouse gases increase the atmospheric absorbed fraction f of longwave radiation from 0.77 to $f' = 0.78$. Calculate the radiative forcing induced by increased greenhouse gases!

Step 1: Calculate temperature in the unperturbed system:

$$F_S/4 = F_S A/4 + (1-f)\sigma T_E^4 + f\sigma T_a^4 = [2f\sigma T_a^4 = f\sigma T_E^4] = F_S A/4 + (1-f/2)\sigma T_E^4$$

$$T_E = [(1-A)F_S/(4\sigma(1-f/2))]^{0.25} = 290 \text{ K}$$

Step 2: Freeze the temperature and calculate the radiative forcing:

$$\Delta F = F_{in} - F_{out} = F_S/4 - F_S A/4 - (1-f'/2)\sigma T_E^4$$

$$\text{Equilibrium in the unperturbed system } \Rightarrow F_S/4 = F_S A/4 + (1-f/2)\sigma T_E^4 \Rightarrow$$

$$\Delta F = F_S A/4 + (1-f/2)\sigma T_E^4 - F_S A/4 - (1-f'/2)\sigma T_E^4 =$$

$$= \frac{1}{2}(f' - f)\sigma T_E^4 = \frac{1}{2}\Delta f\sigma T_E^4 = 2.0 \text{ W/m}^2$$

Climate Sensitivity

- How does ΔF relate to a temperature change when feedbacks are neglected?

Radiative Forcing (greenhouse gases):

$$\begin{aligned} \Delta F &= F_S(1-A)/4 - (1 - \frac{f+\Delta f}{2})\sigma T_0^4 = \\ &= (1 - \frac{f}{2})\sigma T_0^4 - (1 - \frac{f+\Delta f}{2})\sigma T_0^4 = \frac{\Delta f}{2}\sigma T_0^4 \end{aligned} \quad (1)$$

Assume a new temperature equilibrium:

$$\frac{F_S(1-A)}{4} = (1 - \frac{f}{2})\sigma T_0^4 = (1 - \frac{f+\Delta f}{2})\sigma (T_0 + \Delta T_0)^4 \quad (2)$$

Small perturbations:

$$(T_0 + \Delta T_0)^4 \approx T_0^4 + 4T_0^3\Delta T_0 \quad (3)$$

Combine (2) and (3) and neglect second order terms:

$$\Delta T_0 = \frac{T_0\Delta f}{8(1 - \frac{f}{2})} \quad (4)$$

from (1) and (4):

$$\Delta T_0 = \lambda\Delta F;$$

where:

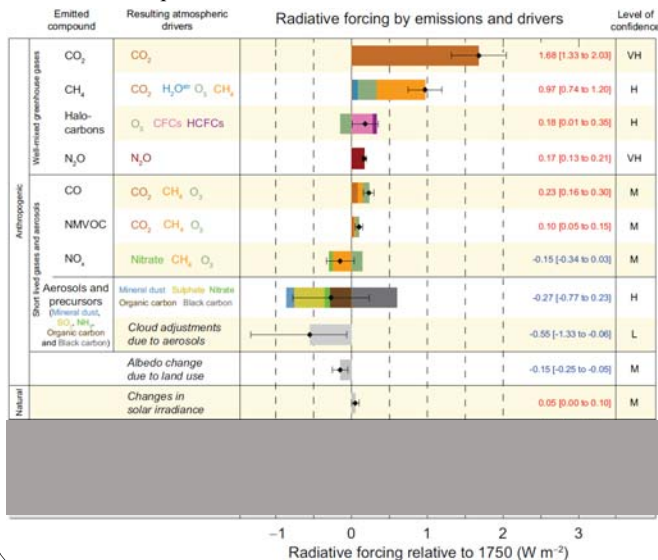
$$\lambda = \frac{1}{4(1 - \frac{f}{2})\sigma T_0^3} = 0.3 \text{ K/(W/m}^2)$$

$$(T_0 = 288 \text{ K}; f = 0.77)$$

λ = climate sensitivity parameter

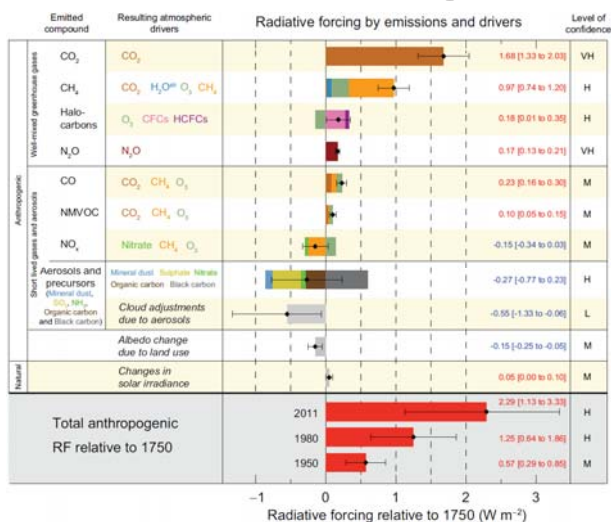
Radiative Forcings of Pollutants

UN climate panel (IPCC):



- Radiative forcing
 - Expresses change in radiative properties or potential climate change
 - Difficult to translate to temperature change due to feedbacks
- The radiative forcing by greenhouse gases (GHG) quantitatively known
- Larger uncertainties in the direct aerosol effect
- The indirect aerosol effect poorly known quantitatively

Radiative Forcings of Pollutants



- Anthropogenic ΔF : +2.3 W/m²
- Climate sensitivity factor (λ) – neglecting feedbacks:
 - $\Delta T = \lambda \Delta F = 0.3 \times 2.3 = 0.7$ K
 - ΔT observed last 100 y: 0.8 K
- Large uncertainty 1.1 – 3.3 W/m²
 - Mainly from aerosols
- => Large uncertainties in the climate sensitivity
- => Aerosol forcings have the potential to mask the warming from greenhouse gases
- Might delay detection of dangerous climate change

We have significantly changed the radiative properties
 - Difficult to estimate temperature change due to feedbacks